

SMOS Instrument Performance and Calibration after 3 Years in Orbit

M. Martín-Neira⁽¹⁾, I. Corbella⁽²⁾, F. Torres⁽²⁾, J. Kainulainen⁽³⁾,
R. Oliva⁽⁴⁾, J. Closa⁽⁵⁾, F. Cabot⁽⁶⁾, R. Castro⁽⁷⁾, J. Barbosa⁽⁷⁾, A. Gutierrez⁽⁷⁾,
E. Anterrieu⁽⁸⁾, J. Tenerelli⁽⁹⁾, F. Martín-Porqueras⁽¹⁰⁾, G. Buenadicha⁽⁴⁾,
S. Delwart⁽¹¹⁾, R. Crapolicchio⁽¹¹⁾

⁽¹⁾ *European Space Agency, ESTEC, Noordwijk (Netherlands)*

⁽²⁾ *Polytechnic University of Catalonia, Barcelona (Spain)*

⁽³⁾ *Aalto University School of Electrical Engineering, Helsinki (Finland)*

⁽⁴⁾ *European Space Agency, ESAC, Villanueva de la Cañada (Spain)*

⁽⁵⁾ *EADS-CASA Espacio, Madrid (Spain)*

⁽⁶⁾ *CESBIO, Toulouse (France)*

⁽⁷⁾ *DEIMOS, Lisbon (Portugal)*

⁽⁸⁾ *IRAP, Toulouse (France)*

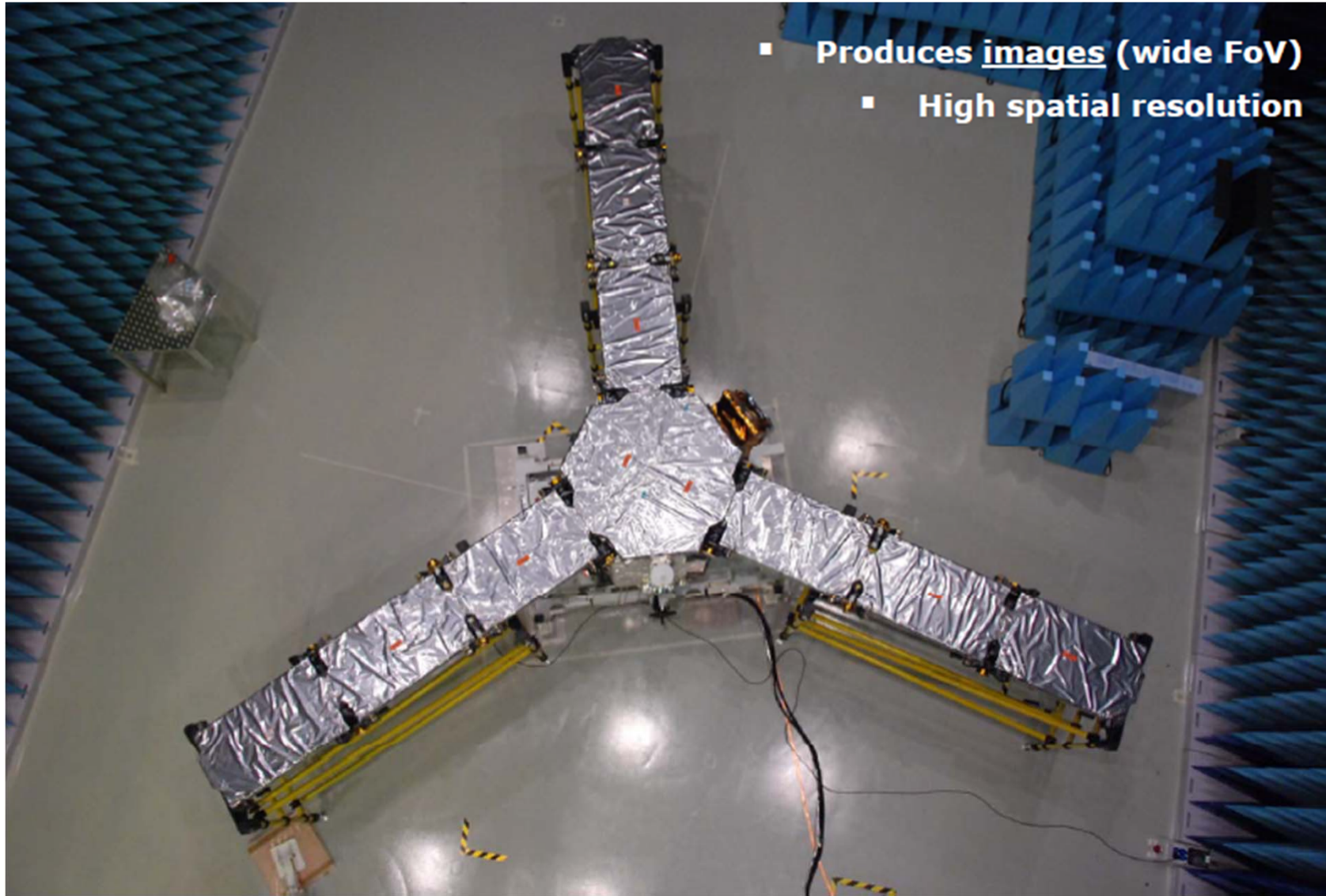
⁽⁹⁾ *CLS, Brest (France)*

⁽¹⁰⁾ *IDEAS, ESAC, Villanueva de la Cañada (Spain)*

⁽¹¹⁾ *European Space Agency, ESRIN, Frascati (Italy)*



MIRAS: the SMOS Payload



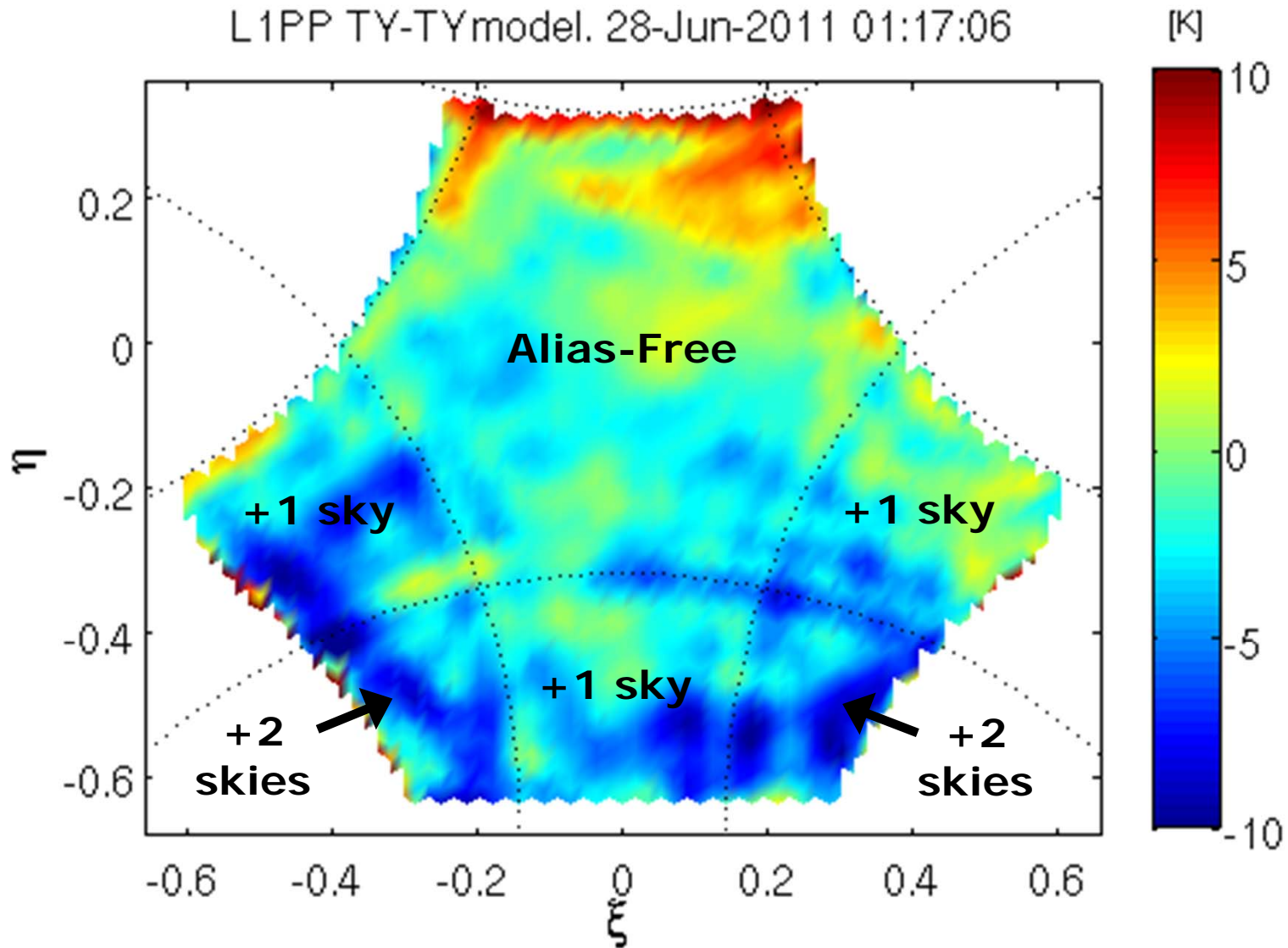
- Produces images (wide FoV)
- High spatial resolution



ALIAS-FREE and EXTENDED FIELDS OF VIEW



L1PP TY-TYmodel. 28-Jun-2011 01:17:06

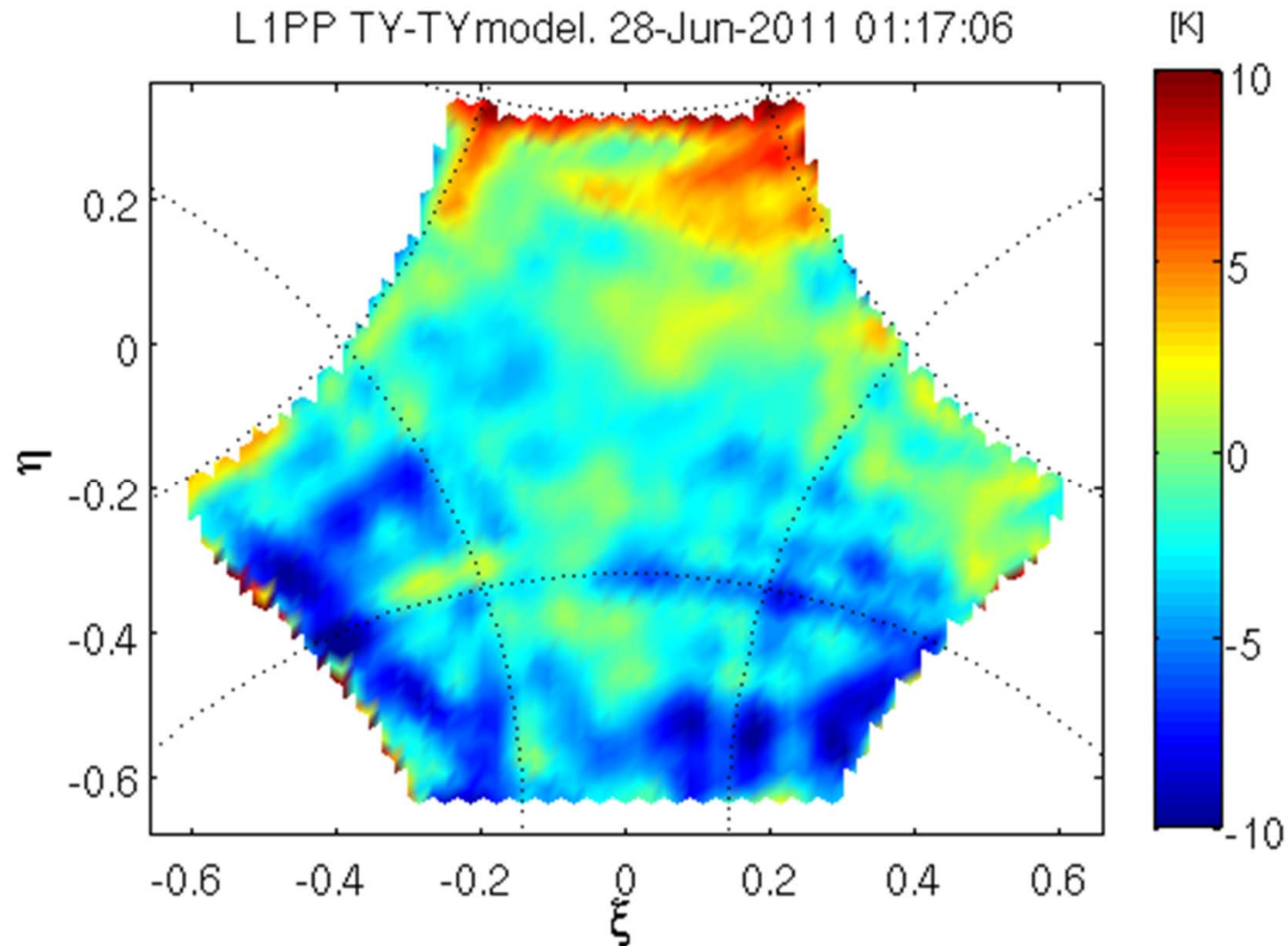


Spatial Biases:

- Ripple
- Slope
- Borders
- Sun tails

Temporal Drifts:

- Orbital
- Seasonal
- Yearly





SPATIAL BIASES: Cold Sky Visibilities



– Cold Sky Visibilities

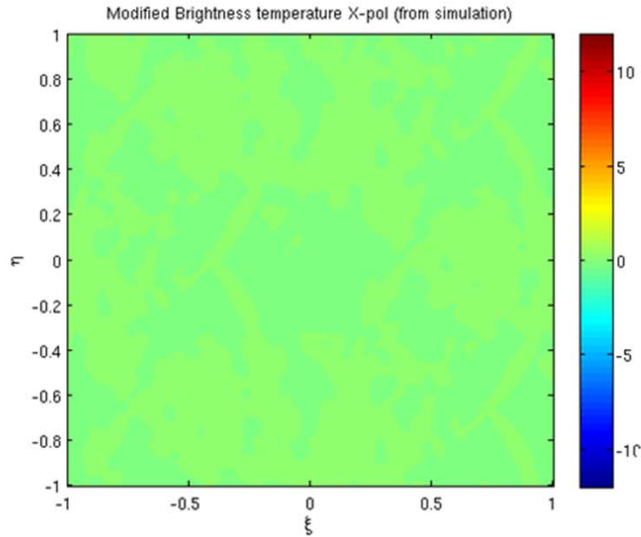
- **improved matching between expected and measured values** by antenna pattern study:
 - use of **average pattern** across 3 frequencies
 - use of **back lobes**
 - **replace pattern of hinge elements** by their neighbours'
 - use of cross-polar (marginal impact)
 - **independent checks** carried out
- remaining discrepancy probably caused by different reasons:
 - **antenna range limitations**
(increased coupling, noise floor, limited reflectivity...)
 - **set-up differences** between ground and flight near hinges
(confirm by electromagnetic simulations)
 - purely image processing issues as **grid type and resolution**



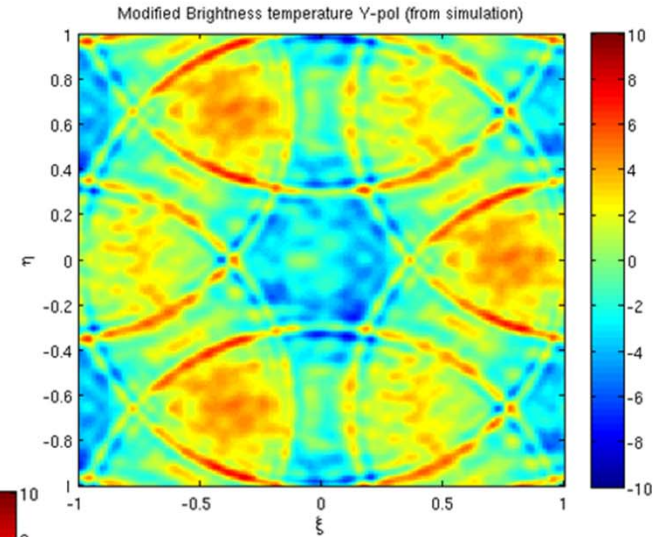
SPATIAL BIASES: Cold Sky Images and the Corbella Equation



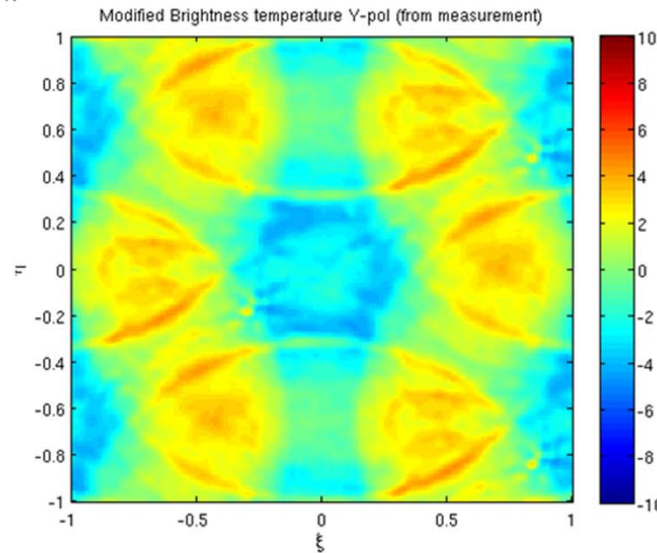
RADIO-ASTRONOMY



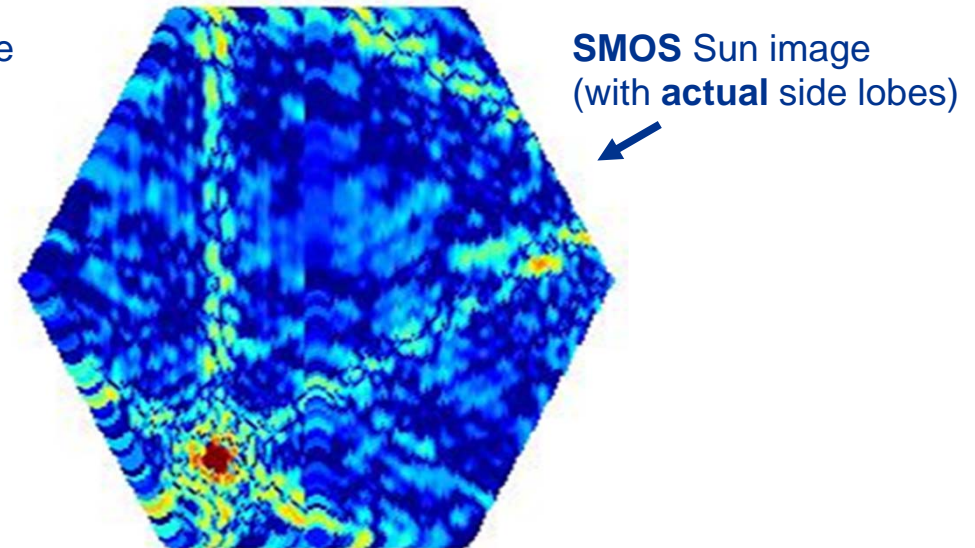
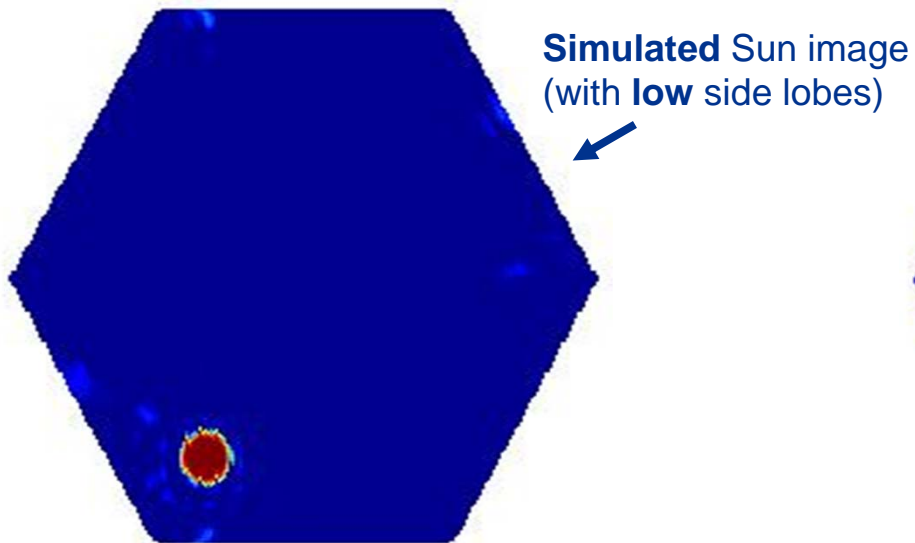
CORBELLA



SMOS MEASUREMENTS



- **Spatial Ripples** are now better understood; they are caused by
 - **side-lobes**
 - **antenna pattern errors**
- **Side Lobes** is the dominant contributor to
 - **spatial ripples**
 - **land-sea** and **ice-sea** contamination
 - **Sun** and **RFI tails** spreading through the image



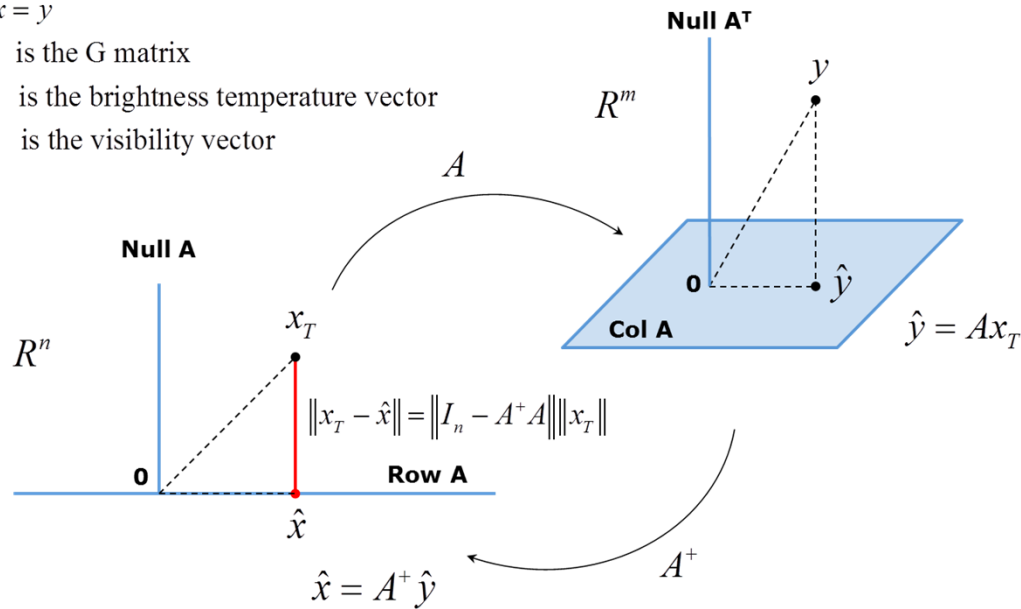
- **Spatial Ripples** are difficult to correct (even for a known instrument)

$$Ax = y$$

A is the G matrix

x is the brightness temperature vector

y is the visibility vector



Kelvin (1 sigma)	SMOS MEASUREMENTS	PERFECTLY KNOWN INSTRUMENT	ANTENNA and CALIBRATION ERRORS
T_x	2.14	1.68	1.33
T_y	2.41	1.25	2.06

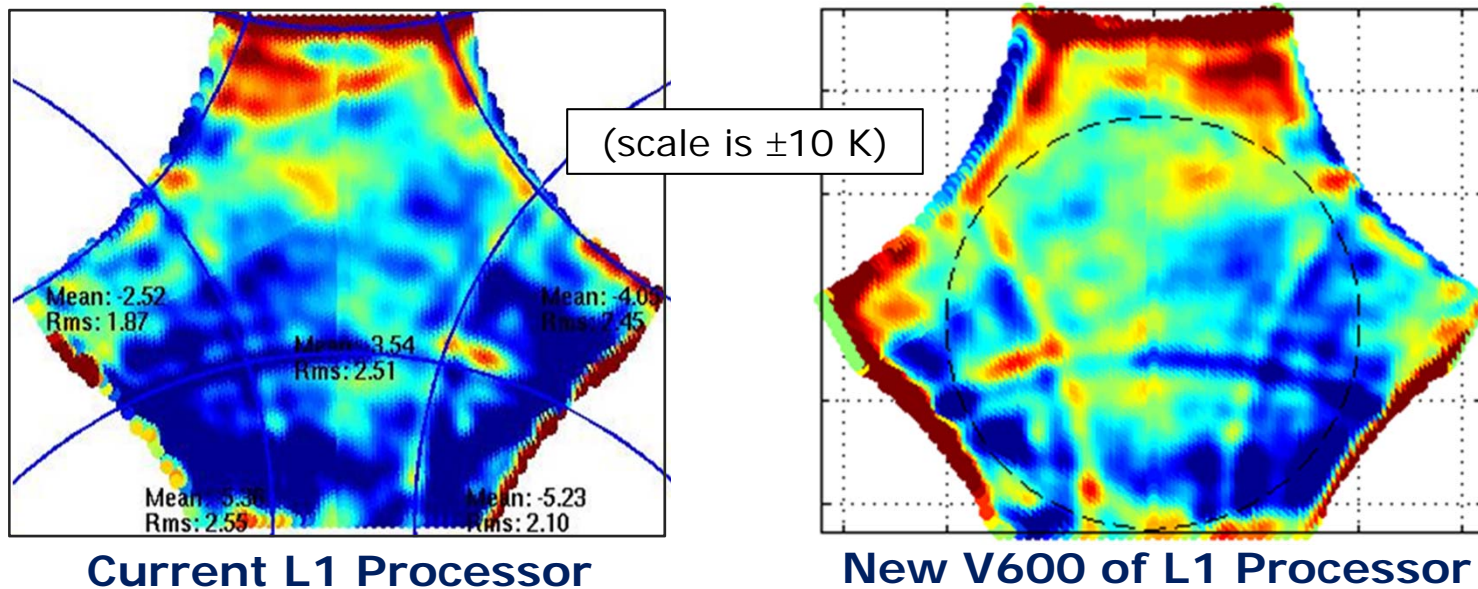
- the **FTT** (Flat Target Transformation) is the best initial correction
- the **OTT** (Ocean Target Transformation) is currently needed for SSS
- tests of a **multiplicative mask** have shown limited results so far
- an **OCT** (Ocean Calibration Target) technique (variation of **OTT**) is under test
- a **review** of the **image reconstruction** started in parallel

– Negative Trend at Low Incidence

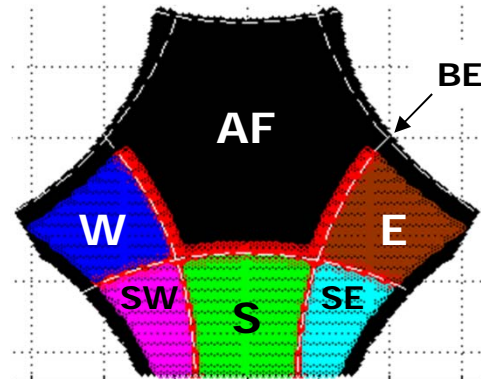
- dependent on L1b image reconstruction approach
- the negative trend has been reduced in the new V600 of the L1 processor

– Borders (belt/suspenders)

- dependent on grid type and resolution
- trade-off between spatial ripple and belt/suspenders in L1b approach



Field of View partition to evaluate metrics of the different Level-1B processors



Spatial Ripple	DPGS	V600 (Model)
X	1.61 K	1.52 K
Y	2.22 K	1.97 K

X-pol	AF	W	SW	S	SE	E	BS
DPGS	0.10 ± 1.30	0.70 ± 1.30	0.20 ± 1.60	-1.30 ± 1.80	0.00 ± 1.70	-0.30 ± 1.30	0.20 ± 2.10
Basic Approach	1.10 ± 1.49	-0.64 ± 1.88	-0.85 ± 3.21	-0.66 ± 1.82	0.32 ± 2.05	1.19 ± 2.81	-0.37 ± 2.05
Model Approach ← best	1.02 ± 1.14	-0.09 ± 0.85	1.15 ± 2.09	0.95 ± 1.30	0.96 ± 1.73	0.45 ± 1.36	-0.88 ± 1.82
FTR Approach	1.08 ± 1.18	-0.17 ± 1.15	-0.08 ± 2.11	0.53 ± 1.27	0.97 ± 1.83	1.78 ± 1.05	1.02 ± 1.82

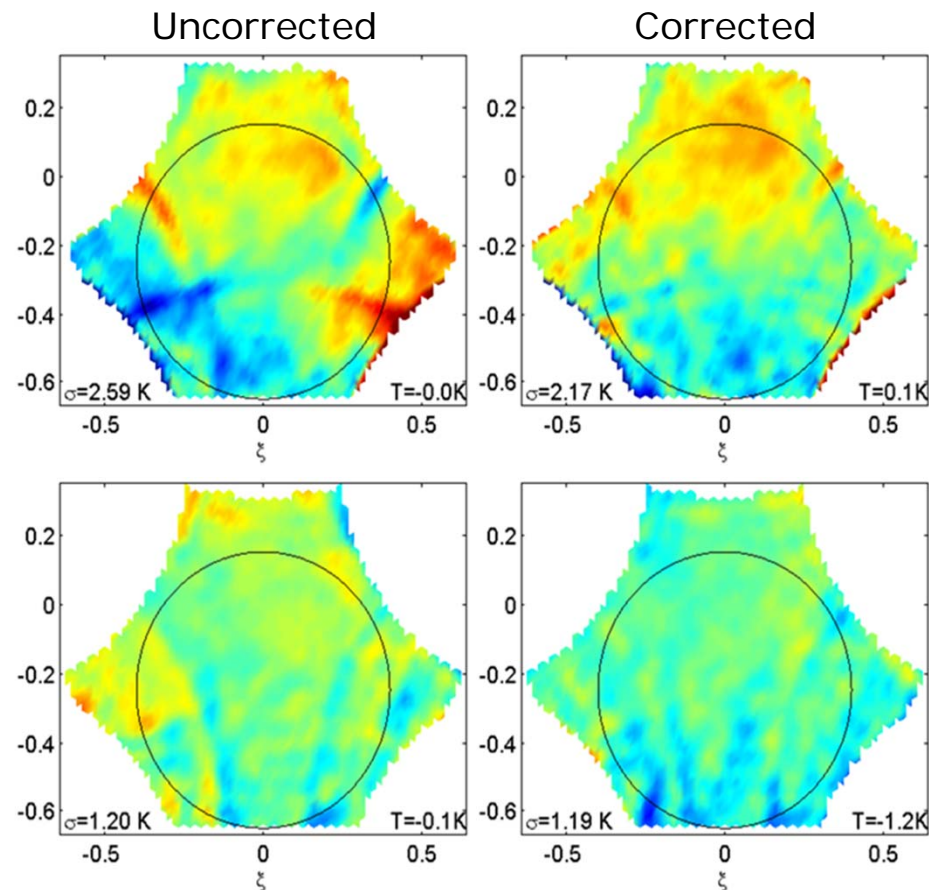
Y-pol	AF	W	SW	S	SE	E	BS
DPGS	-1.10 ± 1.70	-2.50 ± 1.90	-5.30 ± 2.60	-3.30 ± 2.10	-5.20 ± 2.10	-4.10 ± 2.50	-3.50 ± 2.50
Basic Approach	-0.04 ± 1.43	0.87 ± 2.25	-1.45 ± 2.49	1.53 ± 2.32	-1.43 ± 2.62	-0.97 ± 2.20	0.38 ± 2.48
Model Approach ← best	-0.31 ± 1.48	-0.77 ± 1.71	-2.10 ± 2.13	-1.60 ± 1.58	-4.32 ± 2.36	-2.16 ± 2.11	-3.39 ± 2.25
FTR Approach	-0.16 ± 1.40	-0.96 ± 2.15	-2.98 ± 2.14	-1.38 ± 1.75	-4.08 ± 2.24	-1.99 ± 1.54	-1.83 ± 2.43

- The new Level-1 Processor (V600) with the **Model Approach** will bring:
 - **reduced slopes, borders and ripple** (improvement mostly in Y-pol)

– **Slopes** in T3 and T4

- due to **wrong sign** in the horizontal cross-polar antenna pattern
- corrected in the new V600 of L1 processor; hence slopes are removed

Evaluation of
Residual Errors
in T3 and T4
(scale is ± 10 K)

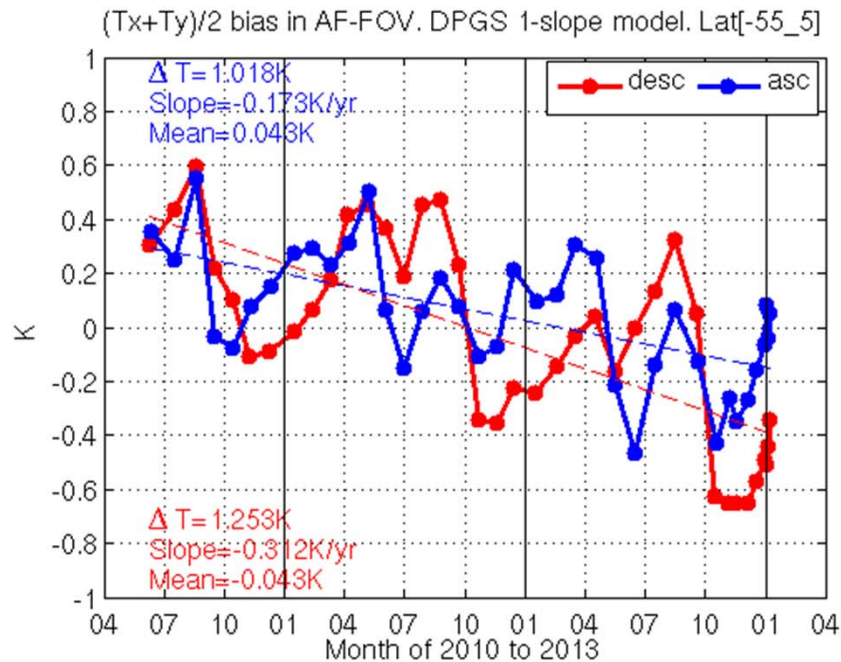


$T3 \equiv \text{Re} \{T_{xy}\}$

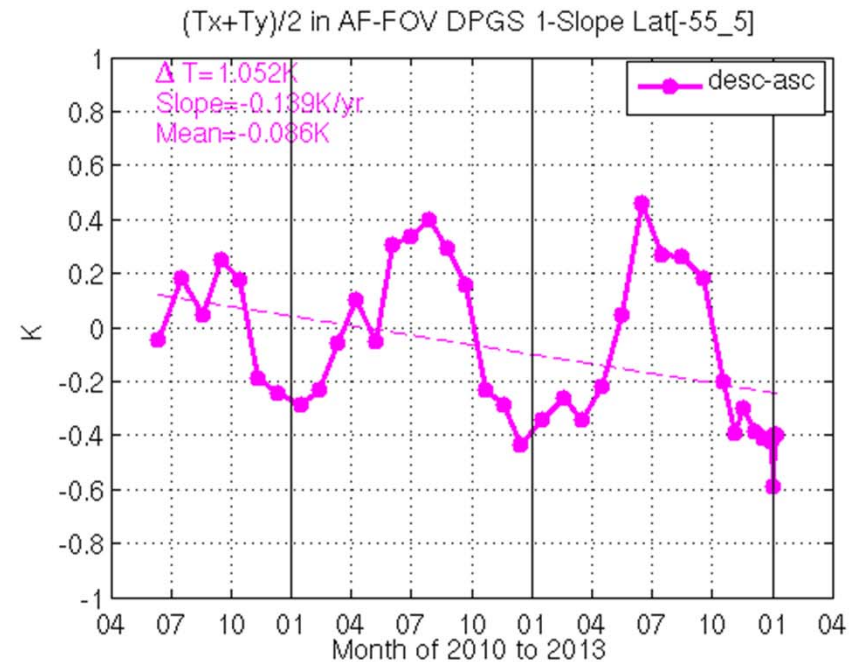
$T4 \equiv \text{Im} \{T_{xy}\}$

Brightness Temperature Residuals over Ocean

Ascending and Descending Orbits Separately



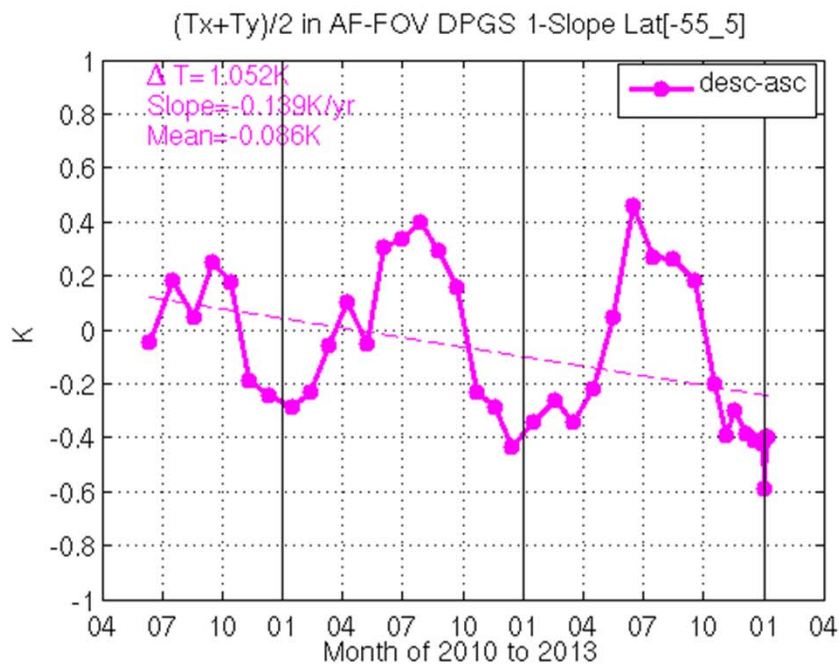
Descending minus Ascending



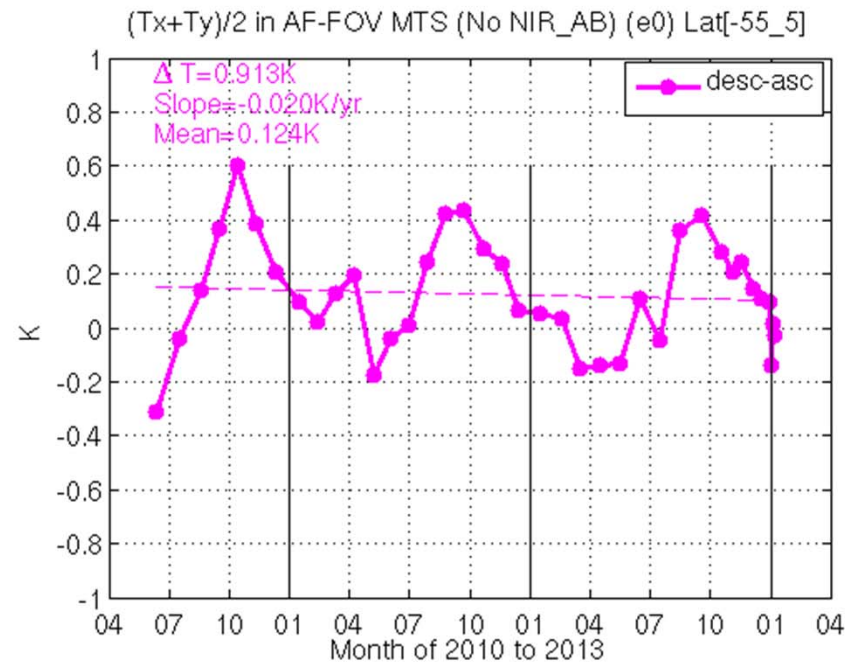
Orbital (peak-to-peak)	1.05 K
Seasonal (peak-to-peak)	1.0 K
Yearly (slope in worst polarisation)	-0.31 K/year

Brightness Temperature Residuals over Ocean

Current Level-1 Processor



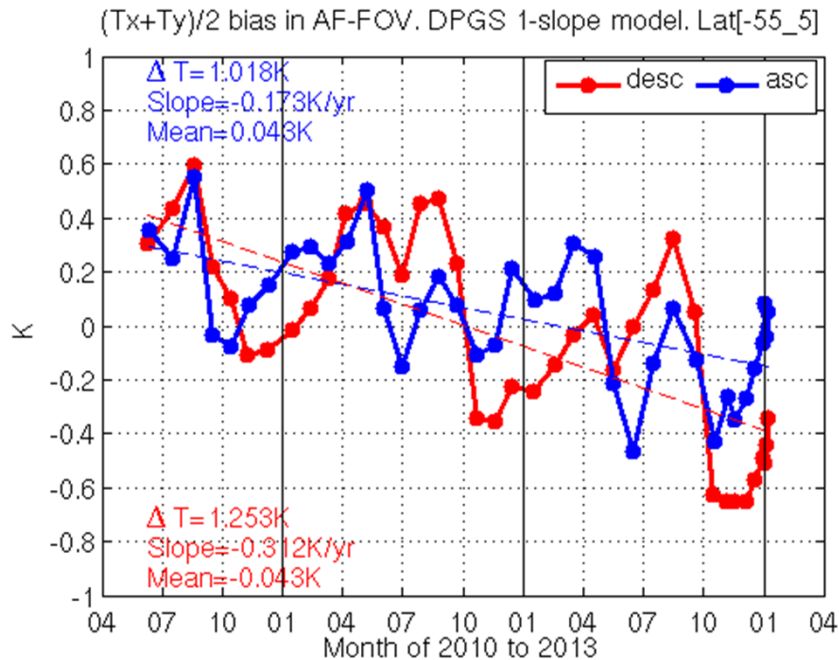
New V600 of Level-1 Processor



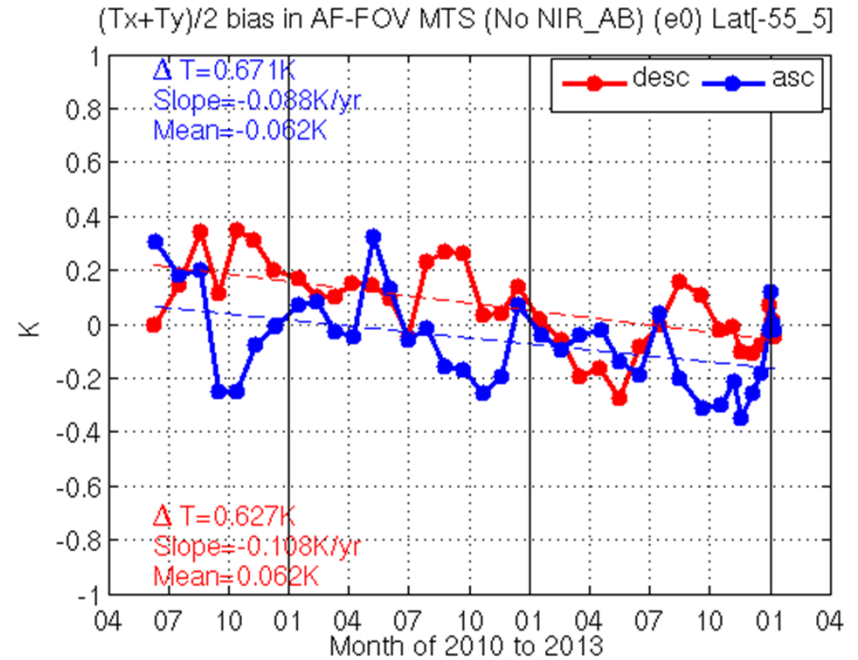
	Current L1	New V600 L1
Orbital (peak-to-peak)	1.05 K	0.91 K

Brightness Temperature Residuals over Ocean

Current Level-1 Processor

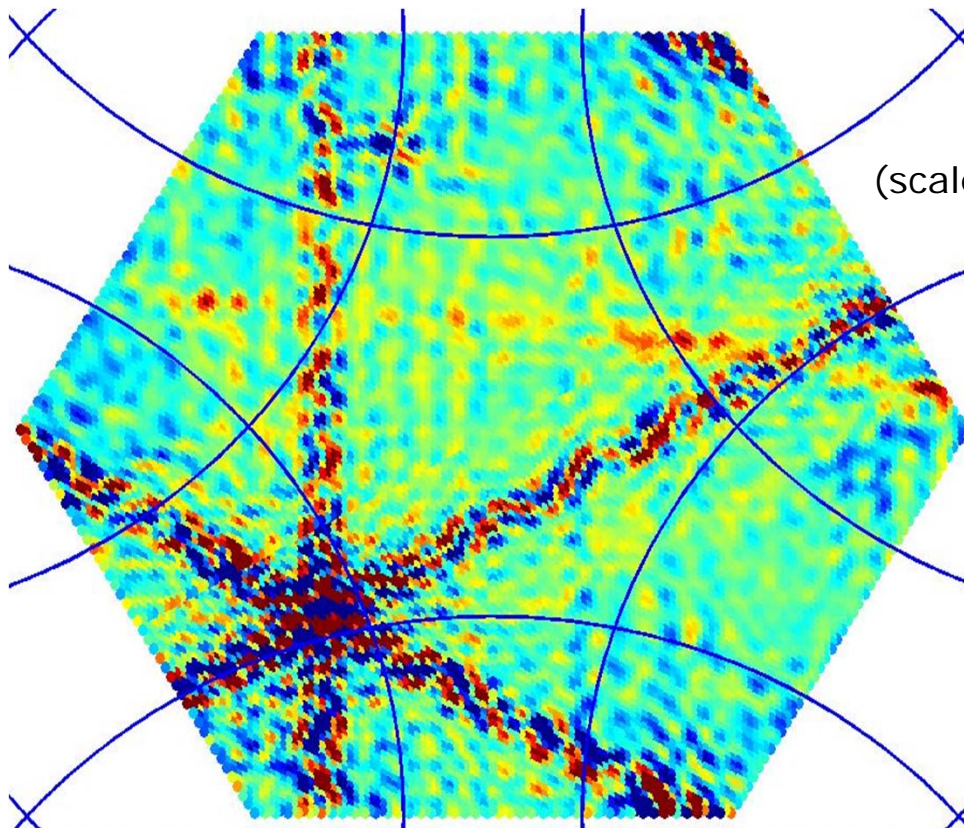


New V600 of Level-1 Processor

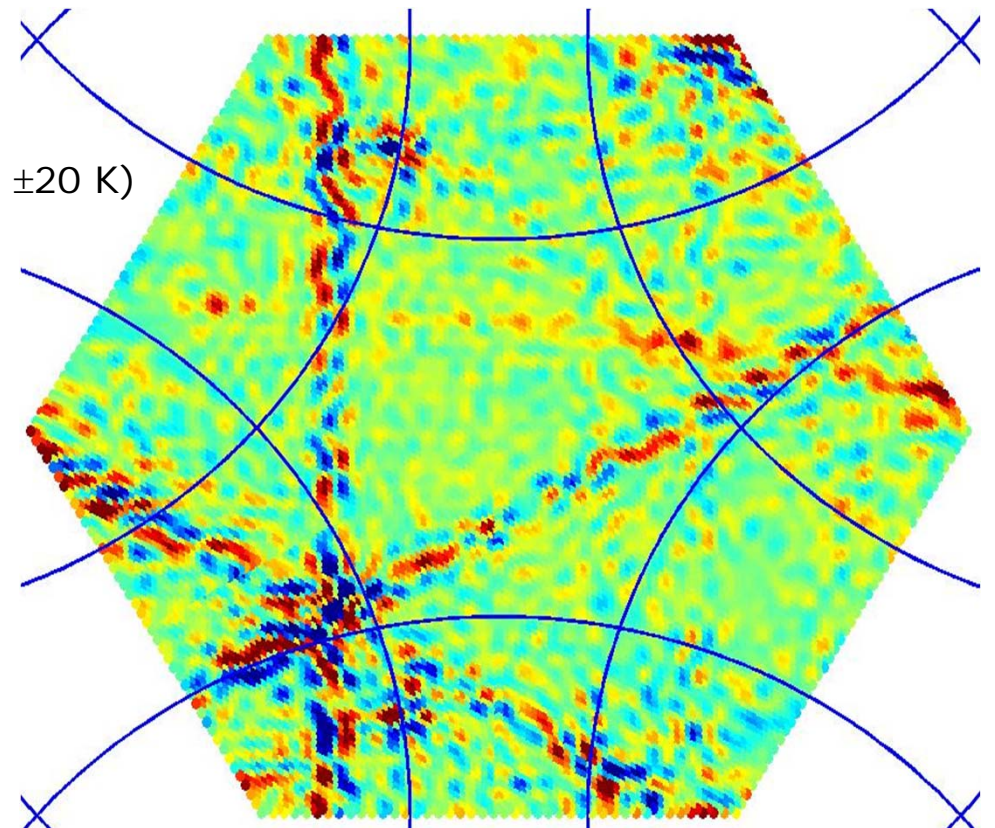


	Current L1	New V600 L1
Seasonal (peak-to-peak)	1.0 K	0.6 K
Yearly (slope in worst polarisation)	-0.31 K/year	-0.11 K/year

Current Sun Correction



New Sun Correction (under test)



(scale is ± 20 K)

- improvement in the range between **20%** to **50%**



CONCLUSIONS



- Corbella equation successfully verified
- Important progress in the understanding of spatial ripples in the images
- Software inconsistencies corrected, resulting in significant improvements
- Optimised Level-1B processor (image reconstruction)
- Reduced slopes, borders and ripples with the new V600 L1 processor
- Instrument stability improved with the new V600 L1 processor
- Full polarimetric data processing with the new V600 L1 processor